COST-STSM-ES1207 Proposal Brewer AOD calibration and retrieval in the UV-B

Izaña Atmopheric Research Centre - PMOD / WRC

WG meeting – Database and processing, Santa Cruz de Tenerife, Spain, 27th – 28th January 2015 Main goals:

- Continuing and completing the development of a standard methodology to transfer aerosol optical depth (AOD) calibration factors between Brewer spectrophotometers, using as a reference a standard instrument
- Define a proposal for a uniform Brewer AOD algorithm

This will contribute to standardize the Brewer aerosols measurements within the scope of the Eubrewnet network.

- Development of a series of algorithms to process AOD absolute calibration (Langley method) for the RBCC-E triad members, including analysis and calculation of correction factors for non-linearity effect on Brewer neutral density filters.
- Establishment of a common algorithm applicable to the Eubrewnet network to compute AOD using a data set of total ozone measurements (DS measurements).
- A comparison of the calculated AOD by Brewer with co-located Cimel data and a UV-PFR.

- Analysis of the Brewer results and assessment of the long-term stability of the Brewer. Seven years of spectral solar measurements will be analyzed (2008-2014).
- Development of a standard methodology to transfer the absolute calibration of direct spectral irradiance measurements from the RBCC-E travelling standard to instruments within the European Brewer network.
- Application / validation of the developed methodology by use of Brewer dataset corresponding to the last RBCC-E Brewer intercomparison (El Arenosillo 2013).

Aerosol Optical Depth Retrieval I

The optical depth of the atmosphere can be divided into several components, ascribed to Rayleigh scattering, aerosols and gaseous absorption:

Aerosol Optical Depth Equation

$$\tau_{\alpha,i} = -\frac{1}{\mu_{\alpha}} \left\{ \ln \left[\frac{l_i(\lambda)}{l_o(\lambda)} \right] -2.6858 \times 10^{16} D_{O_3} \sigma_i(T,\lambda) \mu_{O_3} -\frac{p}{p_o} \beta_i(\lambda) \mu_R \right\}$$

Note that we multiply the ozone optical depth by the factor

$$1DU = 2.6858 \times 10^{16}$$
 molecules cm⁻¹

$I(\lambda)$	\equiv	measured spectral irradiance
$I_o(\lambda)$	\equiv	extraterrestrial spectral irradiance
$\sigma_i(T,\lambda)$	\equiv	differential ozone absorption cross section
		per molecule at temperature T
D_{O_3}	\equiv	total ozone content [DU]
*µ _{O3}	\equiv	ozone airmass factor
p _o	\equiv	mean sea level pressure
р	\equiv	station pressure
$^*\beta_i(\lambda)$	\equiv	Rayleigh scattering coefficient
$^{*}\mu_{R}$	\equiv	Rayleigh airmass factor
$ au_{lpha,i}$	\equiv	aerosol optical depth
$^*\mu_{lpha}$	\equiv	aerosol airmass factor

Aerosol Optical Depth Retrieval III

Open Issues:

- We can neglect gaseous extinction due to *NO*₂, but ... what about *SO*₂? Can RBCC-E provide any *SO*₂ calibration?
- $\bullet\,$ Beer's equation assumes monochromatic irradiance $\Longrightarrow\,$ stray light
- Normal optical depth for Rayleigh as defined in ... Brewer algorithm [Nicolet. M., 1984; Groebner and Meleti, 2004]?, AERONET algorithm [Bodhaine et. al., 1999]?
- Both Brewer and AERONET use the same approximation for air mass factor calculation [Kasten and Young, 1989]
- Should we normalize our calibration factors to 1 U.A.?
- Correcting for internal polarization
- We have to deal with non-neutrality effects on Brewer *Neutral Density* filters

Aerosol calibration is done similar to the ozone one; only it is done for each slit separately rather than the ratios. The basic idea would be, from

$$\ln I_i(\lambda) = \ln I_{i,o}(\lambda) - [\tau_{O_3,i}\mu_{O_3} + \tau_{R,i}\mu_R + \tau_{\alpha,i}\mu_\alpha]$$

to derive the instrument's individual $I_{i,o}(\lambda)$ by imposing $\tau_{\alpha,i}^{instrument} = \tau_{\alpha,i}^{reference}$ for near-simultaneous measurements.

It is important to have the instrument properly characterized (both instrumental and wavelength) as well as ozone-calibrated. Transfer and checking of AOD calibration constants would be desirable as part of ozone calibration

Langley Extrapolation method. We need:

- Direct monochromatic irradiance
- Extraterrestrial irradiance needs to be stable during the measurements included in the Langley plot
- Same as above, for atmospheric conditions
- A wide range of air masses is needed for the Langley plot method ⇒ low-latitude stations

The Izaña Observatory, being a subtropical high-altitude station, allows reliable AOD calibration factors calculation which will be later transferred to the European Brewer network